

# Comparison of the Physics of Wideband/Off-Axis Beams

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Workshop on Long Baseline Neutrino Experiments

March 6-7, 2006

Fermilab

# Outline

- Why off-axis?
- Why wide band?
- $\text{NO}\nu A$  with two detectors
- Wide band beam with one detector
- Comparison
- Summary & Conclusion

# Why off-axis?

The off-axis technology is appealing because

- simple tuning of beam energy
- narrow beam – concentrates the events around the oscillation maximum and allows to do a “counting” experiment
- no high energy tail – high energy neutrinos produce lots of NC events which tend to be reconstructed at low energies
- low background – somewhat reduced  $\nu_e$  contamination

# Why not off-axis?

The off-axis technology has intrinsic limitations

- narrow beam – concentrates the events around the oscillation maximum and reduces to do a “counting” experiment
- background –  $\nu_e$  contamination

Being a counting experiment implies that absolute event numbers are important, thus it is very demanding in terms of systematics. It also means that one can measure only two numbers  $n_\nu$  and  $n_{\bar{\nu}}$ . Virtually impossible to resolve the degeneracies.

# Why off-axis?

The solution to the 'only two numbers' problem is to put a second detector at a different location.

A different location either means a different off-axis angle hence a different energy or a different baseline.

This can result in a different  $L/E$  and thus allows to move into the second oscillation maximum. Where the CP and matter effects are very different.

Or one choose a location with the same  $L/E$  but a very different  $L$  and thus a very different magnitude of matter effects.

see Olga's talk

# Why wide band?

One may consider a wide band beam because

- higher energy (not always an advantage) – longer baseline, more matter effects
- higher on-axis flux
- broad spectrum – many values of  $L/E$  at the same time
- energy information to fight systematics

# Why not wide band?

Wide band beams were 'abandoned' because

- high energy – long baseline for the first maximum reduces flux
- high energy tail – NC feed down, puts stringent demands on the detector
- broad spectrum only useful if the energy resolution is sufficient

This puts the emphasis on the detector side: large mass to compensate distance, good energy resolution and NC rejection

# What do we learn from that?

Just on general grounds, it is not possible to say which approachs works better. To tackle that question a full simulation is required, since the answer depends on many details: energy resolution, NC background, beam power, available baselines, detector technolog, money ...

In the remainder of this talk I try to take what was available to me to approach that goal – I didn't get too close, though.



Some of the following results  
are very preliminary!

# Analysis

Oscillation parameters and errors:

$$\Delta m_{21}^2 = 8 \times 10^{-5} \text{ eV}^2 \pm 10\% \quad \theta_{12} = 0.55 \pm 10\%$$

$$\Delta m_{31}^2 = 2.5 \times 10^{-3} \text{ eV}^2 \pm 10\% \quad \theta_{23} = \pi/4 \pm 10\%$$

Full oscillation analysis including disappearance channels, energy information, systematics, matter density error of 5% with GLoBES.

# The three questions

We want to learn three things from an advanced neutrino experiment

- $\theta_{13} \neq 0$  – if it shouldn't have been found
- $\text{sgn}\Delta m_{31}^2$  – so called mass hierarchy
- $\delta$  – is CP violated in the lepton sector?

Therefore I will use these indicators

- $\theta_{13}$  discovery potential – exclusion of  $\theta_{13} = 0$
- $\text{sgn}\Delta m_{31}^2$ -discovery for normal hierarchy –  
assuming  $\Delta m_{31}^2 > 0$  exclusion of  $\Delta m_{31}^2 < 0$
- CP violation – exclusion of CP conserving values  
 $\delta = 0$  or  $\pi$

# Acknowledgments

The results were obtained in collaboration with: V. Barger, M. Bishai, M. Dierckxsens, M. Diwan, C. Lewis, D. Marfatia, B. Viren and W. Winter

Special thanks to M. Messier who provided the NuMI fluxes on very short notice!

Computing was provided by GLOW and CMS Tier-2 facilities sponsored by the NSF

# NO $\nu$ A + 2nd detector

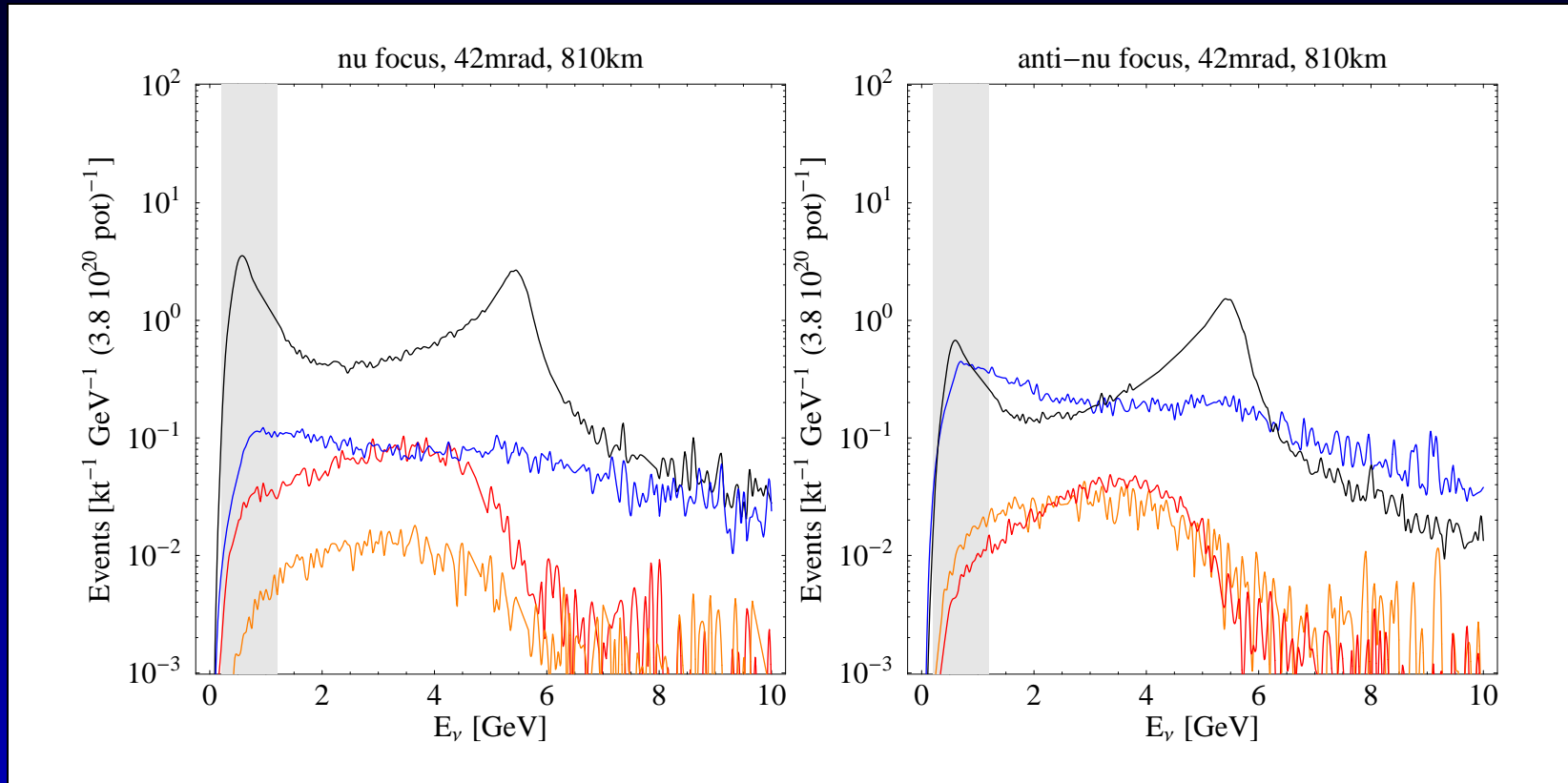
There have been two different ideas on the (US) market:

- 2nd detector at 710 km and 30 km off-axis (42 mrad) – second oscillation maximum  
NO $\nu$ A proposal, 2005
- 2nd detector at 200 km and 8.4 km off-axis (42 mrad) – first oscillation maximum  
O. Mena Requejo, S. Palomares-Ruiz and S. Pascoli 2005

In both cases a 50 kt water Cherenkov detector á la T2K is among the considered options.

Both scenarios assume a FNAL proton driver and 6 years  $\nu$  and 6 years  $\bar{\nu}$  with NO $\nu$ A and 3 years  $\nu$  and 3 years  $\bar{\nu}$  for the 2nd detector.

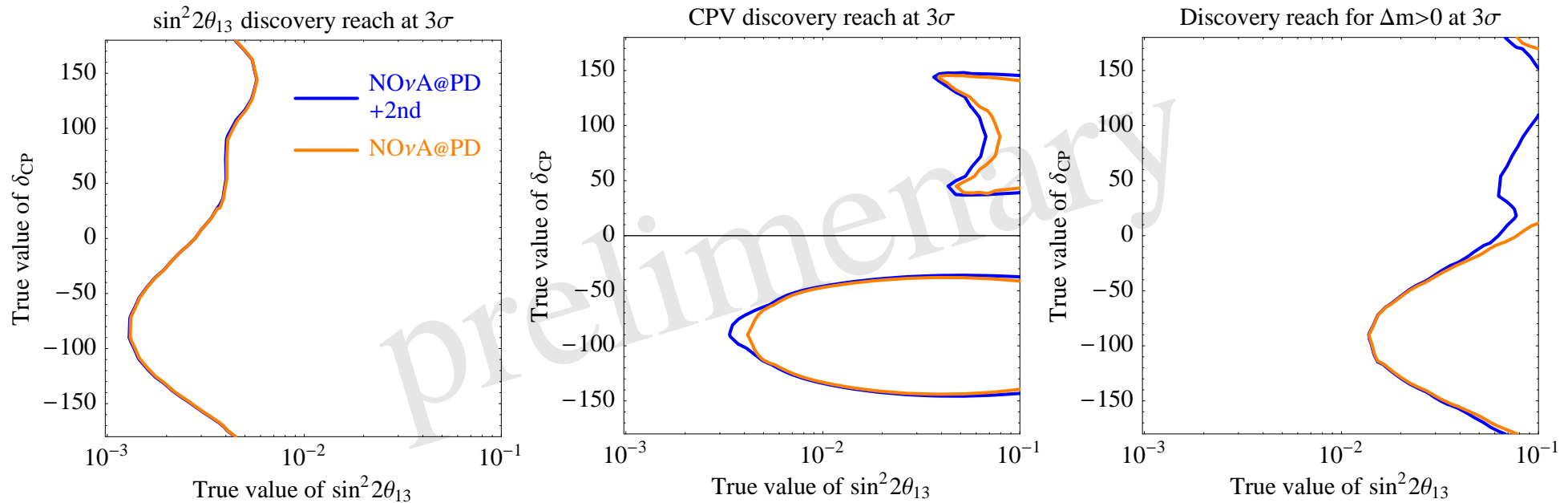
# Beam at 42mrad



- What happens with the second peak?
- $\nu$  background to  $\bar{\nu}$  signal very large
- Only gray shaded region considered here

see B. Flemming's talk

# NO $\nu$ A + 2nd detector



- problems due to  $\pi$ -transit for  $\sin \delta > 0$
- water cherenkov is not optimal
- Super-NO $\nu$ A performs similar

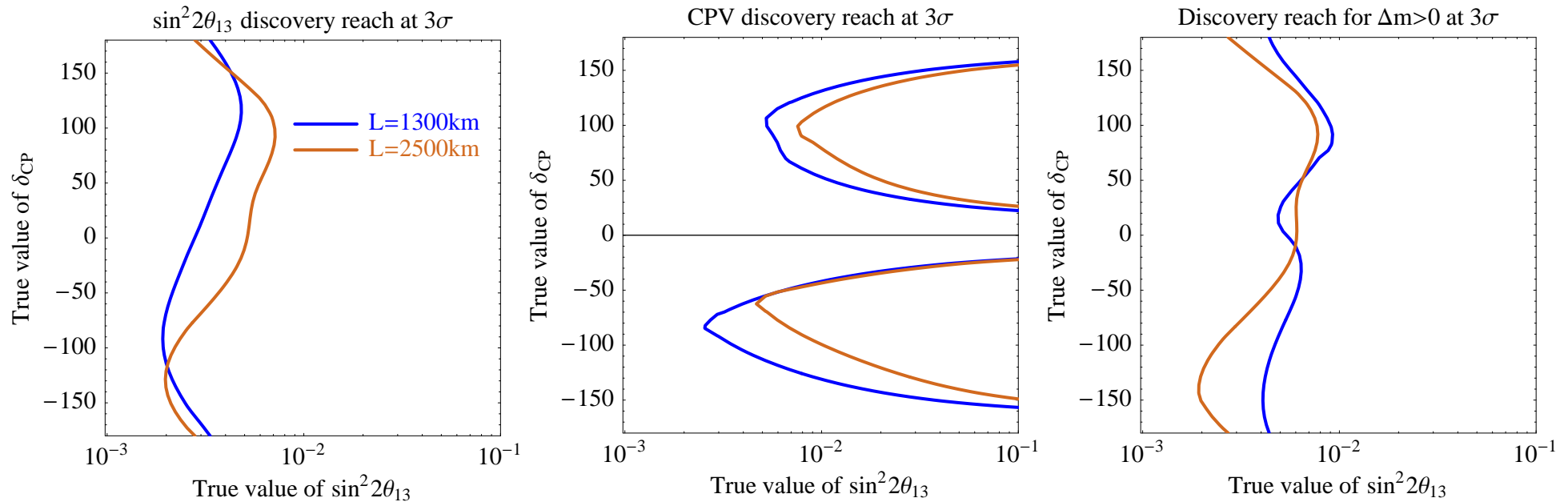
# Wide band beam

- protons with  $E = 28 \text{ GeV}$  and  $P = 1 \text{ MW}$
- 500 kt water Cherenkov detector
- $\pi^0$  suppression verified by Super-K MC  
see Yanagisawa's talk
- $5 \times 10^7 \text{ s}$  neutrino running
- $5 \times 10^7 \text{ s}$  anti-neutrino running
- 10% uncertainty on the background
- $L = 1300 \text{ km}$  or  $L = 2500 \text{ km}$

With the FNAL proton driver this corresponds to 10 years with a 125 kt detector.

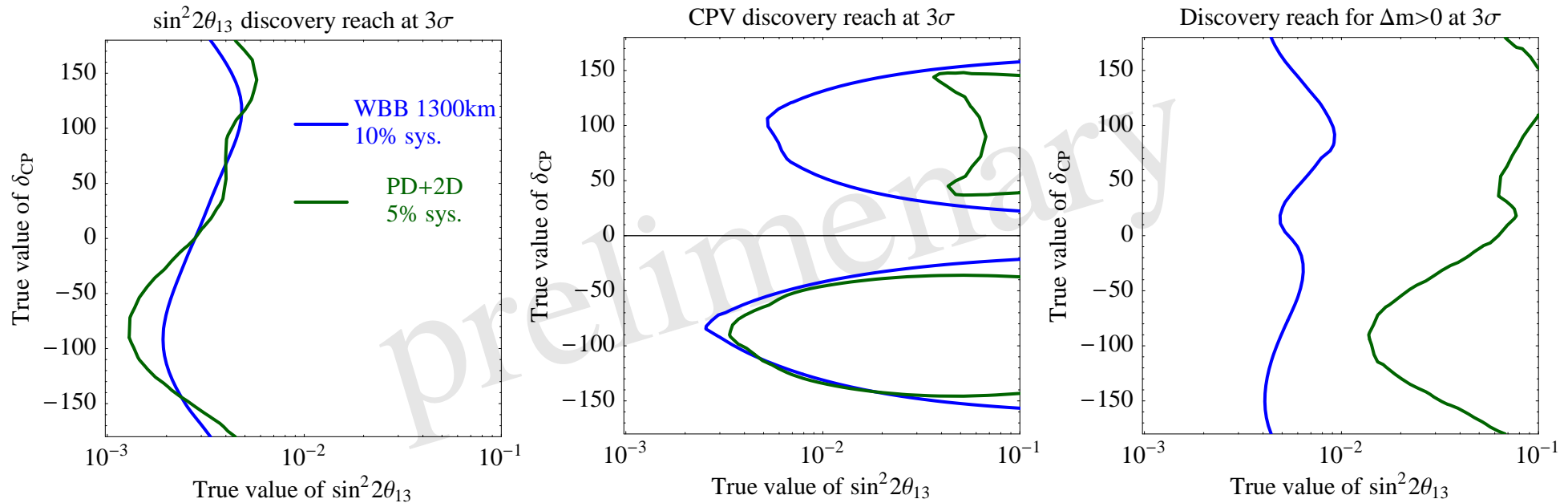


# Wide band beam



- very good resolution of the mass hierarchy
- **no** problems due to  $\pi$ -transit for  $\sin \delta > 0$
- Baseline choice is not critical

# Summary



How would that picture look like with

- Liquid Argon
- 2nd peak in the OA spectrum

# Open issues

- Detector performance is crucial  $\Rightarrow$  need quantitative understanding of the different technologies
- Systematics are important, esp. for OA beams
- How does the US effort compare to e.g. Japan
- ...